

Virunga Volcanoes Supersite Biennial Report: 2022- 2023

1. Abstract

For 2022 and 2023, the activity at Nyiragongo volcano has almost been at background level after the May 2021 eruption. The effusive activity that resumed inside the Nyiragongo main crater in September 2021 began with a sustained seismic activity and ground deformation to the southern flank of the volcano, has shown regular activity with lava presence in the crater captured several times by satellite when weather conditions allow. The degassing is permanent, in fact the two DOAS stations installed around Nyiragongo volcano (one in Rusayo and the other in Sake) indicate that the plume is permanent with 2943 tons/day as maximum daily flux for 2023 and a 300 kt as cumulative SO2 emissions for 2023, from the DOAS that measures only during the presence of sunlight. It has been challenging to access Nyiragongo summit, the May 2021 eruptive site and the lower parts of Nyiragongo volcano (e.g. for measuring temperature and gas in fractures) because of the ongoing war in the volcanic field, which has even prevented the UN helicopter to overflowing the volcanic field in support to the monitoring activities.

Unlike Nyiragongo volcano, Nyamulagira volcano has been characterized by an intense activity inside its large caldera (~2,2x2 km wide), and as for Nyiragongo, Nyamulagira was not accessible for ground-based measurement or helicopter overflowing because of the security issues. The satellitebased monitoring has played an important role in responding to several unrests recorded during 2023 to 2023, especially in May, June and November 2023 when lava overflowed Nyamulagira's caldera rim. Few local scientists are involved in the processing of the EO data, raison why there is still need to continue improving local capacities to process the EO data and for they interpretation. Discussions have been initiated for the capacities exchange between local and international members of the Supersite.

Considering the fact that gases dissolved in the deep waters of Lake Kivu represent a real hazard for the millions of the populations living in its watershed especially those in the cities of Goma (DRC) and Gisenyi (Rwanda), efforts have been made for the studding the stability of this lake and how could a volcanic eruption and/or the movement of the western branch of the rift where the lake lies could affect its stability. Hence, studies on the imaging of the details of the crustal and the upper mantle structure along the Kivu Rift zone particularly the lake Kivu basin as well on the geochemical monitoring of Lake Kivu are ongoing.

Finally, several local scientists had the opportunity to participate in abroad trainings related to volcano monitoring, as well as volcanic and climatic hazards assessment and mitigation. The Goma Volcano Observatory has also received a donation of 5 GNSS stations from the GEO-GNSL, under the Virunga Supersite umbrella, to improve local infrastructures for a better ground-based collection.

2. Scientists/science teams

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Scientists/science teams issues

 The involvement of local scientists in the processing and interpretation of data, both the EO and ground-based has made a step forward through their trainings abroad but there still much to improve. We shall continue with hybrid trainings of local scientists to further improve their capacity.

3. In situ data

**Published data are freely accessible; the Goma Volcano Observatory may be contacted for guidance in case of issues*

***Archived unpublished data or recently collected data are available on request, and based on the data policy for data sharing and management among the Virunga Supersite scientific community.*

In situ data issues

Some in situ data that were collected in partnership with international collaborators may still not accessible to other scientists, despite the Data Policy that was proposed by the Supersite and approved by the GVO's staff, after they have amended it (the data policy is available here, http://geo-gsnl.org/wp-

content/Documents/Supersites/Virunga/History/Virunga%20Supersite_Data%20Policy.pdf). In this data policy, it is stipulated that the "ground-based data and results should be made available for full and open access after publication or in any case within 2 years ("embargo period") since their collection/generation". In any case, the GVO authorities and that of other local institutions should be contacted to get the requirements as per accessing in situ data.

4. Satellite data

<In the table below please list all satellite data types available for the Supersite>

5. Research results

5.1. In situ Data

5.1.1. Seismology : imaging the details of the crustal and the upper mantle structure along the Kivu Rift zone

1. Introduction

The Kivu rift formed in one of the highest elevation regions of the Western Rift Valley of Africa (WRA) hosts Africa's two most active volcanoes: Nyiragongo, and Nyamuragira and other dormant volcanoes and eruptive vents lie under Lake Kivu and its flanks. The Lake Kivu Basin constitutes a transition zone of the WRA from the predominant north-westerly orientation of the Rusizi– Tanganyika Rift Basin to the more southwest to the northeast trend of the Lake Kivu, Virunga, and Rutsuru–Lake Edouard Rift Basins (Ebinger, 1989a). The NE-SW faults dominate the Lake Kivu region. They belong to the Mesoproterozoic Karagwe-Ankole Belt and are overprinted, to the west of Lake Kivu, by NNE-SSW faults of the Neoproterozoic Itombwe Synclinorium which reoriented the rift direction (Villeneuve, 1987; Villeneuve and Chorowicz, 2004). This zone has experienced moderate to large earthquake (e.g The 24 October Kalehe earthquake Mw6.2; the 03 February 2008 Bukavu-Cyangugu earthquake Mw6.0; the 07 August 2015 Katana earthquake Mw5.8). These earthquakes claimed lives, more people were injured, and caused some houses to collapse (Mavonga, 2007; Mavonga, 2009; D'Oreye et al. 2011; Tuluka et al., 2020).

Studies of focal mechanisms of earthquakes occurring in the Lake Kivu region display normal faulting. Bamulezi et al., 2017 identified seismogenic faults on the Western side of the Kivu Rift associated with two distinct large seismotectonic zones. The first coincides with the northern part in which faults follow NNE-SSW orientation and the second comprises the southern part including the Rusizi river until the northern tip of the Lake Tanganyika basin with a general N–S trend, as opposed to the NE–SW trend of the Albertine-Rwenzori segment.

Earthquake hazard assessment of the Democratic Republic of Congo (DRC) Rift zone and adjacent areas including the Lake Kivu region was carried out by Mavonga and Durrheim, 2009 and Tuluka et al., 2020. Their hazard map did not take account of local geology and seismicity at a small scale due to the sparse distribution of seismic stations. Therefore, earthquakes foci were located with large uncertainty. The accuracy of earthquake location is largely dependent on the local velocity structure (Chiarabba et al., 2000).

 Due to this deficiency, to fill this gap, and in order to reduce uncertainty in the location of earthquakes, we have installed additional broadband seismic stations in and around Lake Kivu with Rwanda partners. At this first stage, in the DRC side, two new seismic stations (Kalehe, KALEH; Nyabibwe, NYAB) were built and the old station (Bulengo, BULE) was repaired. These three stations

were equipped with new broadband sensors and digitizers. At the second stage, five broadband seismic stations have been or will be installed on the Rwanda side to fill the gap of coverage of seismic stations in the Lake Kivu area.

Previous study on crustal structure beneath two broadband seismic stations in the Virunga area (North of Lake Kivu, see Mavonga, 2010), determined a crust-mantle transition zone beneath the area sampled by the stations (Kunene, KNN) and (Kibumba, KBB) to be at a depth from about 30 to 41 km. In the Tanganyika rift, south of the lake Kivu region, the crustal thickness ranged from 31.6 to 39 km (Hodgson et al., 2017). Regional seismological studies reported crustal thickness estimates of 30 to 45 km in continental and global models of crustal structure (Mooney et al., 1998; Tugume et al., 2012; Laske et al., 2013). The objective of this work is to take advantage of teleseismic earthquakes waveforms recorded by seven seismic stations located in and around Lake Kivu to bring crustal P-wave velocity constraints to complement hazards studies in an area were earthquakes already damaged cities. At this first stage, seven seismological stations in D.R. Congo maintained by the "Centre de Recherche en Sciences Naturelles (CRSN)" (Lwiro, LWI and Idjwi, IDJ) and the Goma Volcano Observatory (GVO) (Goma, GOM; Bobandana, BOBN; Kalehe, KALEH; Bulengo, BULE and Nyabibwe, NYAB) were used in this study (Figure 1).

All seismic data was acquired at sample rates of 50 Hz and transmitted in real-time at Goma station base. 148 teleseismic events located between 30° and 95° from the stations shown in the figure 1 with magnitude greater or equal to 5.5 were recorded for the period from January 2016 to March 2020. We chose 80 earthquakes with a good signal-to-noise ratio which were used in the analysis. The locations of hypocenters (latitude, longitude, depth, magnitude, and region) of these teleseismic events were obtained from the USGS (Table 1). The located earthquakes and these stations are shown in Figure 2.

 In this report, crustal thickness (Moho), Poisson ratio and P-wave velocities profile for seven broadband seismological stations was estimated in the Kivu Rift in the Eastern Democratic Republic of Congo using receiver function technique. P-wave receiver functions were modelled using H-k stacking analysis including error analysis through bootstrap, variance and phase weighted stacking. The first order velocity profile was obtained by inversion of receiver functions using singular-value decomposition method (SVD).

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Figure 1. Seismic stations of KIVUSNet used in this work, including those acquired through the Project LKMP. These stations are represented by a black triangle.

1°0.00 101°0.00 Legend Stations Earthquakes used in this study $0.5.5$ $• 5.6$ $• 5.7$ $• 5.8$ ö 5.9 6 \bullet 6.1 ö 6.2 a. -00.00 6.3 6.4 6.5 \bullet ۰ 6.6 6.7 ٠ 6.8 ö $• 6.9$ $•7.1$ $•7.3$ $0.7.5$ World Map 250 500 km $\pmb{0}$ $1°0.00'$ 101°0.00'

Figure 2. Distribution of earthquake epicenters used to calculate P receiver functions. The events have a magnitude larger than or equal to 5.5 with an epicentral distance between 30° and 95° from stations GOM, BLG, BOBN, NYAB, KALEH, IDJ, and LWI. These stations are represented by a triangle.

2. Results

2.1. Crustal Thickness and Vp /Vs Ratio Estimates

Using the "hkstack" program which executes an H-k stacking of receiver functions [Zhu and Kanamori, 2000], the corresponding crustal thickness h, and V_p/V_s at these 7 stations are listed in Table 2. Figure 3 shows a sample of the H-k stacking results at GOM station, where the bestestimated H and k values are indicated by the center of the 1 σ error ellipse. It also shows the receiver functions at the corresponding station along with the predicted arrival times of Ps, multiples PpPs, and PpSs+PsPs using the best-estimated H and k values.

 Table 2. Average values of Crustal Thickness H and Vp/Vs Ratio

Figure 3. A sample of the H-k stacking results at station GOM where the best estimated H and k values are indicated by the center of the 1 σ error ellipse.

2.1. I

nversion results of average receiver function at GOM, BULE, BOBN, KALEH, NYAB, LWI and IDJ

Using software ''manyinv" (Ammon, 1997), we achieved the final inversion models for the average receiver function obtained from G1, G2, G3, and G4 quadrants for GOM, BULE, BOBN, KALEH, NYAB, LWI and IDJ. Due to the non-uniqueness in the inversion of receiver functions, eight inversions were executed for each run with one initial model provided in Mavonga (2010). We sorted out the acceptable solutions in each inversion and selected those which fit about 80% or more of the signal power. The model produces one or more synthetic waveforms that compare well with the observed average receiver function at GOM, BULE, BOBN, KALEH, NYAB, LWI and IDJ. A sample of these synthetic waveforms at GOM, KALEH, IDJ and LWI is shown in Figure 4. The corresponding velocity

model at these last stations is shown in Figure 5. By assuming a Poisson's ratio equal to 0.25, the Svelocity model can be calculated using the relation $V_p/V_s = 1.73$. Final results are shown in Figure 6. The reduced average velocity model is listed in Table 3.

Figure 4. Synthetic waveforms that compare well with the observed average receiver function at A) GOM, B) KALEH. C) LWI, and D) IDJ. The selected waveforms fit about 80% or more of the signal power.

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Figure 5. Velocity model obtained at each station using average receiver functions for all quadrants computed at stations (A): GOM, (B): KALEH, (C): LWI, and (D): IDJ.

Figure 6. (A) Average velocity structure obtained by superimposing all velocity structures obtained at GOM, BULE, BOBN, KALEH, NYAB, IDJ, and LWI, respectively. (B) Kivu Region average structure simplified from the arithmetic mean of velocity at each layer in (A). (C) Simplified model created by combining layers in (B) embedded in alternating layers of high and low velocities and showed only the average velocity, taking layer thickness into account as shown in Table 3.

 Table 3. Average local velocity in the Kivu Rift Region

3. Discussion

3.1. Crustal Thickness and Vp /Vs Ratio

From the work of Zana et al (1992) on gravity data obtained from the North of Lake Kivu Region survey, South of Virunga volcanic complex, it was observed that the negative Bouguer gravity anomalies decrease from East to West. The two-dimensional model of these anomalies indicates that a surface layer of low-density material fills a V-shaped trough structure beneath the Lake Kivu area. The Goma station, used in this study, is located on the eastern edge of the V-shaped trough. The crustal thickness beneath the Goma station varies from 36 to 40.0 km. The lower and upper bound of crustal thickness from West to East varies between 24 to 36 Km and 30 to 40 km, respectively. The crustal thickness values observed below stations BOBN, BULE, GOM, IDJ, KALEH, NYAB, and LWI are in agreement with this density model. It decreases from East to West. Compared with published laboratory data (e.g., Christensen 1975, 1996), the V_p/V_s ratio k can be related to rock types. For common rock types, V_p/V_s is particularly sensitive to composition. Increasing the silica content lowers V_p /V_s and high mafic content increases it (Kern 1982). Low V_p /V_s ratio favors the presence of felsic crust rather than significant volumes of partial melts. Based on Table 3 and considering their upper values, KALEH (V_p /V_s = 173-1.98), NYAB (V_p /V_s s = 1.63-1.95), IDJ (V_p /V_s s =1.69 -1.89) and LWI (V_p/V_s =1,60 - 1.78) sample a crust between felsic and mafic while BOBN (V_p $/V_s$ =1.52-1.70) and GOM (V_p/V_s =1.58-1.68) sample a felsic crust. This high (Vp/Vs > 1.78) may indicate crustal modification by magmatic intrusion hypothesis. Figure 7 shows the Moho piercing points of rays at a depth corresponding to the Moho depth beneath stations KALEH, BOBN, NYAB, IDJ, LWI, BULE, and GOM, respectively, as shown in Table3. The Moho piercing points were computed using IASP91 (Crotwell et al., 2001). This interpretation is supported by the position of these stations located on the Western shoulders of the Kivu sector of the Western rift where a chain of strombolian cones erupted alkaline basalts and basanite during 2-0 Ma. This alkaline volcanism was coeval with transitional silica-saturated basaltic eruptions on the rift floor and subaqueously (in

part) on the bed of Lake Kivu (Kampunzu and Lubala 2012). Observation of piercing points in Figure 7 shows that our study samples also the Virunga area including volcano Nyiragongo and Nyamulagira but there is some deficit in the Rwanda side area.

Figure 7. Moho piercing points of rays at a depth representing the Moho depth beneath stations GOM, IDJ, NYAB, KALEH, BOBN, LWI, and BULE, as shown in Table 3 and represented in the legend by colour delimited by the intervals 0 to 1, 1 to 2, 2 to 3, 3 to 4, 4 to 5, 5 to 6 and 6 to 7, respectively.

4. Conclusion and recommendations

We evaluated the structure of the crust below seismic stations KALEH, BOBN, NYAB, IDJ, LWI, BULE, and GOM by inverting stacks of teleseisms receiver functions. The resulting first order velocity models were presented as the P-wave velocity model. A lateral variation in receiver function with back azimuth was observed. This implies lateral heterogeneities or dipping of layers. The crustmantle transition zone below the Kivu Rift region is approximative at a depth from 24 to 40 km. Based on the plot of piercing points at Moho depths, a deficit in the Rwanda side was observed. Thus, the result obtained in this study can be improved by setting up the five supplementary broadband stations at Rwanda side at Rubavu, Karongi, Waw island, Nyamasheke, and Rusizi.

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5.1.2. Measurement of sulfur dioxide in Nyiragongo volcanic plume

During an eruption, fluids often precede solids and are therefore "warning signs" of a volcanic eruption. This is why volcano observatories, including the Goma Volcano Observatory (GVO) are interested in monitoring the composition of volcanic gases, particularly of open system volcanoes.

Figure 8. Map showing the location of DOAS stations sites (red dotes) around Nyiragongo and Nyamulagira volcanoes, in the north basin of Lake Kivu. The basement image shows the plumes of the two volcanoes, and image that was acquired on February 9, 2015 by the NASA's Earth Observatory Landsat 8, accessed September 28, 2015) are shown Nyiragongo and Nyamulagira permanent plumes pushed west-southwest ward by the regional dominant wind.

Thus, the GVO studies the composition of gases from the Virunga active volcanoes, i.e. Nyamulagira and Nyiragongo, to capture any changes in their composition which would be a sign indicating possible volcanic eruption. Differential Optical Absorption Spectroscopy (DOAS) instrument is therefore used to measure sulphur dioxide $(SO₂)$ amount in volcanic plume. Four stations were deployed around the active volcanoes with one installed in Rusayo, Sake, Kunene and another in Buzi (Figure 9). The stations experienced several technical difficulties until only the Rusayo station

operated alone before the May 22, 2021 eruption of the Nyiragongo volcano. In March 2022, two new stations were installed in Sake and Rusayo to replace those which had become dilapidated and obsolete. Currently, the network stations are therefore insufficient at the time when the technician continues to struggle for the proper functioning of Sake station. Note that only Rusayo station is functional and the data used in this report are those produced by this station. In the framework of the present study, it is planned to buy for GVO two DOAS instrument, with each consisting of a UV spectrometer, a telescope, an optical fiber, a motor-controlled scanning device, a self-contained remotely programmable microcomputer, a GPS receiver, and a timer.

Volume of sulphur dioxide emitted by Nyiragongo Volcano in 2023

The collected and processed data since the beginning of 2023 come from the station installed in Rusayo, which was used to produce the below graph:

Figure 9. Chart showing the variation of SO₂ concentration in Nyiragongo volcanic plume as measured at Rusayo site between late January to December 2023 at Rusayo station shown in Figure 9.

The SO2 flux recorded at Rusayo, along with satellite imageries (e.g. Sentinel-2 L2A), testify to a continuous effusive activity inside the crater of the Nyiragongo volcano, and which is confirmed by the visibility above the Nyiragongo crater of the gas plume during the days with good meteorological conductions and a glow of light during the night. The areas without data mostly correspond to days when the weather conditions were very bad and therefore did not allow the DOAS spectrometer to analyse the quantity of $SO₂$ due to the lack of sufficient sunlight. $SO₂$ flows vary from 235.71 tonnes/day to 2943.38 tonnes/day with an average of 872.61 tonnes/day. The vast majority of daily $SO₂$ flows are less than 1000 tonnes/day, which remains modest compared to the $SO₂$ flows already emitted by the Nyiragongo volcano during periods of high activity, in fact values of around 18,000

tonnes /day have already been recorded by DOAS at the Nyiragongo volcano. After the May 2021 eruption of the Nyiragongo volcano, activity resumed inside its crater in September 2021, and it continues to gradually increase.

5.1.3. Monitoring of Lake Kivu in the Kabuno bay

Kabuno Bay is a sub-basin of Lake Kivu located to the northwestern part of the lake (Fig. 4), limited to the north by lava flows from Nyamulagira volcano contrary to the main basin of Lake Kivu which is bordered by Nyiragongo lava flows.

Figure 10. Map showing the sites where water samples are progressively being collected at different depth in all basins of Lake Kivu on DR Congo side except Kabuno Bay and where well as CTD profiles will be made.

This bay has a maximum depth of 140 km contrary to the Main Basin which is 485 m, but contains higher amount of carbon dioxide than corresponding depth of the Main Basin. In fact, high carbon dioxide concentration in Kabuno is found from 15 m depth, while in the Main Basin such concentration is found at about 250 m; in the Main Basin. This exposes Kabuno Bay to be more exposed to the risk of an overturn than the main basin, and conducted the RDC authorities to degassing carbon dioxide. The degassing must be done with care to not disturb the chemistry and biology of the bay, explaining the need to establishing a baseline for Kabuno Bay which consist of collecting water samples and their analysis for major cations and anions, as well as CTD multiparameter probe deployment to make vertical pH, temperature, dissolved oxygen, specific

conductivity and redox potential profiles. We also have conducted similar measurement in the main basin of Lake Kivu.

Preliminary results show that the alkalinity of water, expressed in the form of bicarbonate, increases with depth. In the biozone or surface water (from 0 to 60m, in the main basin), the alkalinity is around 700 mg/L in the main basin (Figure 11) and around 1000 mg/L in Kabuno Bay (Figure 12). In the main basin, from a depth of 300 m to the bottom of the lake, the alkalinity is around 4500 mg/L (Figure 11), while from 40 m in Kabuno Bay the alkalinity is already at beyond 4500 mg/L and reaches values of 8500 mg/L at the bottom (Figure 12).

Figure 11. Vertical profiles of total alkalinity in the main basin of Lake Kivu from samples collected between January and December 2023. Total alkalinity was titrated in the form of bicarbonate (HCO₃) either in the field or in laboratory at the Goma Volcano Observatory (OVG), using diluted 0.1 M HCl and methyl orange as an indicator.

Figure 12. Vertical profiles of total alkalinity in Kabuno Bay from samples collected between January and December 2023. Total alkalinity was titrated in the form of bicarbonate (HCO₃) either in the field or in laboratory at the Goma Volcano Observatory (OVG), using diluted 0.1 M HCl and methyl orange as an indicator.

- The temperature profiles in the main basin (Figure 13 A) and in Kabuno Bay have the same trend (Figure 14 A), apart from the fact that in Kabuno Bay the surface waters are 1°C colder than the surface waters of the large Lake Kivu basin.
- The waters of the two northern basins of Lake Kivu have very high conductivity values and show similar increasing trends as the depth increases (Figure 13 B and Figure 14B), these high conductivities are due to the contributions from hydrothermal vents that enter the lake through underground springs and the dissolution of volcanic rocks in the northern watershed of the lake. Despite their shallow depths (140 m maximum), the waters of Kabuno Bay show very high conductivity compared to the deep waters of the large basin which is 485 m deep.
- Other parameters show similar trends. However, chlorophyll A in the surface waters of the waters of Kabuno Bay is higher than in the large basin, on the other hand, turbidity is much higher in the waters of the large basin than in those of the waters of Kabuno Bay.

Figure 13. Vertical profiles of temperature (A), electrical conductivity (B), pH (C), chlorophyll A (D), turbidity (E), and dissolved oxygen (F) in the main basin of Lake Kivu in 2023.

Figure 14. Vertical profiles of temperature (A), electrical conductivity (B), pH (C), chlorophyll A (D), turbidity (E), and dissolved oxygen (F) in the Kabuno bay, a sub-basin of of Lake Kivu in 2023.

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5.2. Satellite Data

5.2.1. The May 2023 Nyamulagira intra caldera Eruption: Change detection map from COSMO-SkyMed Imagery in the period April 28 to May 30

An analysis of the changes that occurred over Mount Nyamulagira during the 2023 May eruption is undertaken to detect and evaluate the lava flow evolution.

Figure 15. Results of processed CSK intensity images as for 28 April 2023, 13 May 2023, 14 May 2023, 22 May 2023, 28 May 2023

According to https://volcano.si.edu/showreport.cfm?wvar=GVP.WVAR20230517-223020, lava flows began moving into the N and NW parts of the crater beginning on 9 May, towards the low point of the crater rim. The lava continued to erupt from vents during 17-23 May.

For this analysis a time series of Synthetic Aperture Radar (SAR) COSMO-SkyMED (CSK) imagery have been used. The change detection methodology is based on the SAR intensity difference between two images collected before and after the event. Figure 15 shows the processed CSK intensity images. From the images in Figure 15, the evolution of the SAR signal intensity over the period can be appreciated in and around the Nyamulagira crater. In this case, the presence of a lava flow should be represented by a lighter shade of gray (increase of SAR signal backscattering) in the successive intensity image dates.

 A visual comparison between (a) (28 April) and (b) (13 May) shows that changes inside the caldera have occurred, with a general increase of the intensity inside the crater. It can be also noted a movement of the lava flow into the NW part of the crater.

No changes are appreciated between 13 May (b) and 14 May (c).

Instead, large differences are visible between 14 May (c) and 22 May (d) due to the eruption onset. In particular, the lava flow covered the N and SW parts of the crater and it started to flow out in the W flank. In particular, the extent of the new lava coverage towards the N can be appreciated by the reduction in size of the dark area adjacent to the N rim.

The visual comparison between (d) (22 May) and (e) (30 May) shows that the lava flow continues to flow out in the W flank that, even in this case, can be appreciated by a reduction of the dark area close to the W flank.

To better identify these changes, a change detection approach based on the intensity difference is applied to each pair of contiguous CSK images. Results are shown in Figure 15 that confirms the visual intensity analysis previously undertaken. The changes associated with the lava flow are represented by dark pixels. The flank lava flow extension is manually traced and shown as a red polygon.

Figure 16. Change detection maps between a) 28 April and 13 May; b) 13 May and 14 May; and c) 14 May; 22 May; and d) 30 May. The changes associated with the lava flow are represented by dark pixels. The W flank lava flow pattern is evident in c) (22 May) and d) (30 May) and shown by a red polygon in e) and f), respectively.

5.2.2. The October-November 2023 Nyamulagira intra caldera Eruption monitored by satellite-based techniques

a. Change Detection mapping from COSMO-SkyMed imagery intensity signals

This section reports the analysis of the changes that occurred on Mount Nyamuragira during the 2023 October-November eruption, undertaken to detect and evaluate the lava flow evolution.

Figure 17. Processed SAR-CSG intensity images collected in ascending pass between October and November 2023, as for 3 October 2023 (A), 19 October 2023 (B), 4 November 2023 (C), 12 November 2023 (D), 12 November 2023 (E).

For this analysis, a time series of Synthetic Aperture Radar (SAR) COSMO-SkyMed Second Generation (CSG) imagery (Italian Space Agency) collected in ascending and descending passes have been used. The change detection methodology is based on the SAR intensity difference between two images collected before and after the event.

1) Ascending pass analysis

Figure 17 shows the processed CSG intensity images collected in ascending pass. The images in Figure 17 show that the SAR signal intensity evolution over the period can be appreciated inside the Nyamuragira crater. The presence of a lava flow is represented by a lighter shade of gray (because of an increase of SAR signal backscattering) in the successive intensity image dates. In particular:

- the visual comparison between (a) (3 October) and (b) (19 October) shows that changes inside the crater can be appreciated with an increase in the backscattering in (b). The two images appear similar outside the crater.
- The visual comparison between (b) (19 October) and (c) (4 November) shows changes inside the crater while no changes are visible outside the crater. However, it is worth noting that the backscattering signal decreases in (c), confirming that the lava is moving inside the crater.
- The visual comparison between (c) (4 November) and (d) (12 November) shows that even in this case, most of the changes can be appreciated inside the crater with an increase of the backscattering signal in (d).
- The visual comparison between (d) (12 November) and (e) (20 November) shows slight changes inside the crater.

To better identify the changes, a change detection approach based on the intensity difference is applied to the pair of CSG images. Results are shown in Figures 17A, B, C and D for the couple (a)- (b), (b)-(c), (c)-(d), and (d)-(e), respectively, which confirms the visual intensity analysis commented above. Most of the changes are visible inside the crater. In Figures 18A and 18C (Figures 18B and D) we can appreciate an increase (decrease) in the backscattering that is represented by a darker (lighter) area.

Figure 18. Change detection map between 3 October and 19 October 2023 (A), Change detection map between 19 October and 4 November 2023 (B), Change detection map between 4 and 12 November 2023 (C), Change detection map between 12 and 20 November 2023 (D).

2) Descending pass analysis

Figure 19 shows the processed CSG intensity images collected in descending pass.

Figure 19. Processed SAR-CSG intensity images collected in descending pass between October and November 2023, for 1 October 2023 (F), 17 October 2023 (G), 2 November 2023 (H), 10 November 2023 (I), 18 November 2023 (J).

Considering the images reported in Figure 19:

- The visual comparison between (f) (1 October) and (g) (17 October) shows that changes inside the crater can be appreciated with an increase in the backscattering in (g). The two images appear similar outside the crater.
- The visual comparison between (g) (17 October) and (h) (2 November) shows changes inside the crater, while no significant changes are visible outside the crater. The backscattering signal decreases in (h), confirming that, even in this case, the lava is moving inside the crater.

- The visual comparison between (h) (2 November) and (i) (10 November) shows slight changes inside the crater with an increase in the backscattering signal in (i).
- The visual comparison between (i) (10 November) and (j) (19 November) shows no changes.

To better identify any changes, the change detection results are shown in Figures 5, 6, 7, and 8 for the couple (f)-(g), (g)-(h), (h)-(i), and (i)-(j) (of Figure 19), respectively, which confirms the visual intensity analysis discussed above. Most of the significant changes take place inside the crater. In Figure 5 (Figure 6) we can appreciate an increase (decrease) in the backscattering that is represented by a darker (lighter) area. In Figures 7 and 8, slight changes inside the crater can be appreciated

Figure 20. Change detection map between 1 and 17 October 2023 (A), between 17 October and 2 November 2023 (B), between 2 and 10 November 2023 (C), between 10 and 18 November 2023 (D).

b. Lava flow mapping from Sentinel-1 imagery coherence signals

This section reports the analysis performed with SAR images to evaluate lava flow due to the 2023 Mount Nyamuragira volcanic activity. Such analysis consists of the monitoring of the interferometric coherence variations induced by the changes due to the presence of lava. The coherence, indeed, indicates the correlation between two SAR images being sensitive to any changes of the scenario occurring in the time interval between the acquisitions. Then, any changes of the observed scenario

due to the volcanic activity produce variations of the coherence values. It can be expressed as follows:

$$
\gamma = \frac{\left|\sum_i x_i y_i^*\right|}{\sqrt{\sum_i x_i x_i^* \sum_i y_i y_i^*}}
$$

where *xi* and *yi* are the pixels of the first and the second SAR image.

For such analysis several pairs of SAR images acquired by Sentinel-1 missions of the European Space Agency from March to November 2023 have been used (Table 4).

Table 4. Sentinel-1 SAR dataset used for the coherence variations analysis.

Figure 21 shows coherence maps estimated from Sentinel-1 SAR images.

e)	October 20th - November 1st	GEOS 1.40°S 1.41's $1.42^{\circ}3$ 1.43° S 0 500 1000 29.18'E	29.19°E	29.20'E	29.21°E	\blacksquare \blacksquare 0.1 \blacksquare 0.2 10.3 \Box 0.4 \Box 0.5 0.6 \Box 0.7 0.8 \Box 0.9 \blacksquare 29.22°E
f	November 1 st - November 13 th	ICEI 1.40° S 1.41's 1.42's $1.43^{\circ}8$ 0 500 29.18'E	1000 m 29.19°E	29.20°E	29.21'E	\blacksquare \blacksquare 0.1 \blacksquare 0.2 \Box 0.3 \blacksquare 0.4 0.5 0.6 \blacksquare 0.7 \Box 0.8 \Box 0.9 \exists 1 29.22°E

Figure 21. Coherence maps estimated from Sentinel-1 SAR images.

The areas characterized by low values of coherence (black zones in Figure 21) in all the pairs outside the crater represent the areas covered by vegetation which is strongly affected by temporal decorrelation effects. Instead, the low coherence largely observed inside the crater until June is due to the presence of a lava lake characterized by lava movements and the coherence loss in the pair May 17th - May 29th corresponds to a lava flow outside the crater and on the West flank of the

volcano.

No significant changes are observed in October and November. The lava field detected in May (and some other previous lava flows) is characterized by high values of coherence in the images of October and November, meaning that the lava cooled, since there are no changes between different pairs.

c. Ground deformation analysis

Thanks to the support provided by the Argentina Space Agency (CONAE) to the Virunga Supersite, it was possible to obtain two images from the SAOCOM satellite, acquired by a SAR sensor in the L band.

We investigated the ground deformation at Nyamuragira by processing the SAOCOM image pair 20231013-20231106, along the ascending orbit (incidence angle of the sensor equal to 36.24°) and in the STRIPMAP acquisition mode, calculating the interferogram shown in Figure 22.

Figure 22. Wrapped interferogram from SAOCOM ascending data between October 13 and November 6, 2023.

The obtained interferogram unfortunately shows some orbital residual fringes, preventing the observation of possible ground displacement related to the unrest episode. With the aim to retrieve information about the current volcanic episode, we considered the coherence map reported in Figure 23.

Figure 23. Interferometric coherence from SAOCOM ascending data between October 13 and November 6, 2023.

 Figure 24. Zoom over the area of the Nyamuragira volcano - Coherence map.

The small black circle within the crater in Figure 11 identifies coherence loss between the two SAOCOM acquisitions (13 October and 6 November). This can be interpreted as a change of the surface that could be related to lava flow inside the crater. This result agrees with the outcomes from the coherence analysis shown previously in this report.

d. Multispectral satellite analysis

The tasking of Skysat and Prisma satellites was activated over Nyamuragira from 10 November 2023. While the Prisma satellite didn't provide good acquisitions for bad weather forecasts, the Skysat constellation, which consists of about 20 satellites, was able to acquire two scenes on 14 and 19 November (Figure 25). From these images (0.5 m spatial resolution at ground), it is possible to observe the melt portion of the lava lake extending in the south-west part of the crater and covering an area of about $15x10^3$ m² (insets of Figure 25).

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Figure 25. Skysat scenes in false-color RGB (Band 4,2,1) acquired on 19 November (top) and 14 November 2023 (bottom). Insets show the lava lake area.

By comparing the PlanetScope images (3 m spatial resolution) of 7 April 2023 and 22 July 2023, we derive the overflow occurred in May - June 2023. The lava flow outside the caldera measures about $0.9x10^6$ m² with the longest branch reaching 3.4 km (Figure 13).

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Figure 26. PlanetScope scenes in false-color RGB (Band 8,4,2) acquired on 22 July (top) and 7 April 2023 (bottom). On the top image the lava flow emplaced in May - June 2023 is clearly visible to the West of the crater.

e. Conclusions

The change detection analysis from the intensity of the SAR images (COSMO-SkyMed) in both orbits evidenced an increase or decrease of lava flow inside the crater from the beginning of

October to the end of November 2023. These changes are limited to inside the crater of Nyamuragira. The coherence mapping of the SAR images (Sentinel-1) in both orbits doesn't evidence changes inside nor outside the crater between the end of October and November 2023. On the other side, during May - June 2023 the presence of a lava lake inside the crater and coherence loss due to a lava flow outside the crater on the West flank of the volcano was detected. The interferogram from SAR images (SAOCOM) between October 13 and November 6 2023 didn't show a clear deformation pattern, although the interferogram is affected by orbital residual fringes. The interferometric coherence loss in a circle within the crater can be due to lava flow inside the crater, probably related to the June 2023 eruption as obtained through the analysis of the Sentinel-1 data. Finally, multispectral satellite imagery allowed to detect and quantify the lava flow field emplaced outside the caldera in May - June 2023 and showed how the thermal activity is currently confined in the southwest portion of the lava lake.

GROUP ON

5.2.3. Thermal satellite monitoring of Nyamuragira volcano

The thermal activity of the Nyamuragira lava lake was followed using the CL-HOTSAT system (Ganci et al., 2016), through the processing of MSG-SEVIRI and Landsat imagery. In particular, SEVIRI data were exploited to retrieve the radiant heat flux time series, while thermal maps derived from Landsat were used to track the thermal activity.

Figure 27. SEVIRI-derived radiant heat flux time series (in Watt) over Nyamuragira volcano from 15 September 2023 to 1 February 2024

The most intense thermal activity occurred from 3 October to 18 November, with peaks overcoming 1.4 GW on 4, 13, 16 and 31 October and the maximum peak of 1.81 GW reached on 5 November. After that, a clear decrease in the radiant heat flux is observed, with values mostly below 0.2 GW until 1 February. The decrease in radiant heat flux since the end of November, as well as the thermal map obtained from Landsat (Figure 28), doesn't show any significant overflow from the lava lake.

Figure 28. Maps of thermal anomalies derived from the Landsat 8 image (30m spatial resolution) acquired on 28 January 2024 at 20:16 GMT.

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5.2.4. Synthesis of Nyamulagira activity from late 2022 through 2023

Since 2014, an intermittent lava lake inside the pit crater located within the Nyamulagira caldera has been observed (Campion, 2014; Smets et al., 2014, 2015; Coppola et al., 2016). The eruptive activity increased in March 2023 with several lava overflows from the pit crater, with the lava that reshaped the caldera floor. On the evening of 19 May 2023, a glow was observed above Nyamulagira's caldera from the city of Goma, suggesting a mighty activity. The lava overflowed the caldera rim in the following day, moving along the volcano's western flank. In the GSNL - Virunga Supersite initiative framework, I used SAR images from Sentinel-1 and Cosmo-SkyMed satellites provided by the European Space Agency and the Italian Space Agency, respectively, to analyse the ground deformation during this peak activity (March-May 2023) and map the lava flow extent. Radar images are fundamental to studying the activity of this volcano since it is located in a remote area with limited access, and the cloud coverage of the tropical environment limits the usage of satellite optical images.

SAR images have been processed using Gamma software (Wegmuller et al., 1998). I evaluated deformation maps for the ascending and descending orbits of the Sentinel-1 satellite using the stacking technique to reduce the noise. To do this, I processed interferograms with a short temporal baseline (12 days between the two acquisitions). The images were multilooked at 13x2 in range and azimuth directions, respectively, to obtain slightly square pixels with a resolution of ~30x30m² and reduce the noise. The SRTM 1 arcsec has been used to remove the topographic phase component and geocode the final products. The interferograms have been filtered with an adaptive filter (Goldstein and Werner, 1998).

The minimum cost flow unwrapping algorithm has been used to convert the phase to

deformation. A common reference area, considered stable, has been selected northwest of the caldera. A residual orbital plane has been evaluated and removed. All the final deformation maps have been summed to evaluate the total displacement in the covered period (Figure 29).

Since Cosmo SkyMed SAR images have a higher spatial resolution (namely 2x2m²), I used the amplitude images to detect and roughly map the lava flow extent (Figure 30). Indeed, the new lava changes the backscat-tering power of the surface. I processed images acquired before and after May 19, 2023, in ascending and descending orbits. The images were multilooked at 3x3 in range and azimuth to obtain ~6x6m² pixels. This is the best way to reduce the speckle noise on the amplitude images.

Deformation maps

The Sentinel-1 ascending and descending deformation maps obtained by stacking the interferograms show a signal in range decrease (movement toward the satellite - up to ∼3 cm) in a large area southwest of the caldera (Fig. 1). This signal could be consistent with an uplift produced by an inflating deep magma chamber (at ∼30km) slightly shifted southwest of the caldera. The presence of a deep source at Nyamulagira was already introduced to explain some far field displacement (e.g. Wauthier et al., 2013). However, the signal-to-noise ratio for the interferograms is low, and this signal could be a residual noise that the stacking technique, with the simple correction applied, couldn't remove (e.g. Biggs et al., 2007). A proper multitemporal InSAR analysis with the use of a larger number of interferograms with a longer temporal baseline and the use of external data to reduce the atmospheric noise (e.g. GNSS zenith delay or weather models) could improve the results confirming the presence of the uplift. Furthermore, it could provide information on the temporal evolution of the ground deformation that is possibly correlated to the eruptive activity inside the caldera.

Lava flow

The changes in the amplitude of the Cosmo SkyMed images acquired before and after the eruptive episode allowed for roughly mapping the lava flow extent (Fig. 2). The lava overflowed from the pit crater located on the caldera's northern part and started filling the northwestern quadrant. The lava then went over the western edge of the caldera and began to flow along the side of the volcano, covering about 2km.

Figure 30. Cosmo skymed amplitude images before and after the lava overflow event.

The differences between the lava extent in the ascending and descending images are related to several reasons, e.g.:

1) the different acquisition times, possibly the lava spread at the tip of the lava flow between May 20 and May 22.

2) the different looking geometries, particularly the incidence angle, produce different reflections that highlight the shape of the lava flow.

3) the entire caldera floor is coated with fresh lava. Indeed, the backscattering power of a two days old lava is comparable to that of a few months old one. Gamma software has masked some pixels (black areas) since they are affected by the shadowing and layover effects (artefacts related to the topography and the incidence angle of the signal).

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5.2.5. Contribution of the satellite monitoring system MOUNTS to the surveillance of Nyiragongo and Nyamulagira volcanoes (DRC)

1. MOUNTS monitoring system

MOUNTS (Monitoring Unrest from Space, www.mounts-project.com, Valade et al. 2019) is a satellite-based volcano monitoring system, exploiting the diversity of sensors on-board the polar-orbiting Sentinel spacecrafts (ESA Copernicus). The synergistic use of radar (Sentinel-1 SAR), short-wave infrared (Sentinel-2 MSI) and ultraviolet (Sentinel-5P TROPOMI) payloads allow for surveillance of surface deformation, emplacement of volcanic deposits, detection of thermal anomalies, and emission of volcanic $SO₂$ (Figure 31). Deep Learning "plugins" are incorporated to assist specific processing tasks, such as the automatic detection of large deformation (INSAR) and SAR speckle filtering. The system is openly inspired by the MIROVA system (Coppola et al. 2020), whereby volcanic activity is visualized through the systematic display of geo-referenced images and time-series of the above-mentioned parameters at several tens of active volcanoes worldwide.

Figure 31. Screenshot of the MOUNTS web interface (www.mounts-project.com) showing the monitoring of Nyamulagira volcano through Sentinel-5P (SO2 mass in purple), Sentinel-2 (hot pixels in orange, Massimetti et al. 2020), and Sentinel-1 (SAR amplitude in gray, INSAR displacement in red, INSAR coherence in blue).

2. Nyiragongo

The May 2021 eruption of Nyiragongo was followed by a drastic reduction in the $SO₂$ emissions and thermal anomalies, consistent with the drainage of the summit lava lake (see Virunga Supersite Biennal Report 2020-2021). In the days following the onset, Sentinel-1 SAR images showed a fast collapse of the crater, evidenced by an increased size and depth of the bottommost pit crater.

Analysis of the Sentinel-1 SAR images allow recovery of the depths variations of both lava lakes and/or pit crater. This is done by measuring the length of the shadow cast by the crater walls in radar geometry, and multiplying it by the cosine of the radar incidence angle (Wadge et al., 2011). The method has been applied to track lava lake level variations at Nyiragongo (Barrière et

al., 2018) and Erta Ale (Moore et al., 2019), and at Nyamulagira it has been applied by MOUNTS to provide variations of the pit-crater floor (Burgi et al. 2021).

Analysis of Sentinel-1 images shows the collapse which occurred in May 2021 following the lava lake drainage (Figure 32). Interestingly, the analysis also shows that after the collapse the crater has to slowly refill, by approximately 69 m between June-2021 and March-2024. Inspection of Sentinel-2 images (Figure 32, inset) suggests that this progressive crater infilling is related to successive episodes of small intra-crater lava flows.

Figure 32. Analysis of Sentinel-1 products (descending orbit) showing the crater collapse in the days following the 2021 eruption onset, and the progressive crater infilling in the following years. Top-most images represent the Sentinel-1 acquisitions displayed in radar geometry, before the eruption onset (2021-05-22), after the onset (2021- 06-02), and now (2024-02-29). Yellow arrows indicate the shadow cast by the bottom terrace, where a decrease in shadow width indicates crater infilling. Inset images are Sentinel-2 images (SWIR band composition) which suggest that this infilling is caused by successive lava flows inside the crater.

3. Nyamulagira

Much like the analysis presented for Nyiragongo, analysis of Sentinel-1 SAR images allows recovery of the changes in the floor depth of Nyamulagira's pit crater and main caldera. Both Burgi et al. 2021 and MOUNTS's contribution to Virunga Supersite Biennal Report 2020-2021 showed that during the period 2018-Jan to 2020-June, the pit crater floor at Nyamulagira suffered variations, characterized by a phase of fast collapse (fall of the crater floor in May/June 2016), which was both preceded and followed by phases of slow re-filling (rise of the crater floor). Extending this analysis up to 2024 (Figure 33b), we notice that the pit crater floor suffered similar phases of collapse and refill. In particular, a large collapse episode was observed in May-2021 (~120 m drop), and a second slower collapse was observed in Aug/Nov 2023 (~40 m drop). Following the large collapse episode in May-2021, the inner-pit crater slowly started refilling. Once the pit was completely refilled, lava flows started to emplace onto the main caldera floor, causing a progressive infilling of the caldera and a subsequent decrease of caldera outer wall heights (Figure 33a).

Figure 33. Variations of (a) the main caldera floor and (b) the pit crater floor at Nyamulagira during the period 2018/04 to 2024/03, analyzed by MOUNTS from Sentinel-1 images. Top-row images: Sentinel-1 SAR images (ascending orbit 174) in radar geometry; yellow arrows indicate the SAR shadow used to recover depths of both the caldera floor and pit crater floor along the radar Line of Sight (LOS). Top-row right-most image: high-resolution optical image acquired on 2022-03-03 (Google Earth), dashed red line indicates the radar LOS along which the depths were calculated.

The infilling of the main caldera (Figure 33a) eventually facilitated lava flows to overflow outside the caldera walls onto the western slopes of the edifice, as observed from the INSAR coherence map on 2023-05-29 (Figure 34). During this event, the entire caldera floor was recovered in lava flows, as shown by the decrease in the mean coherence value (Figure 33b, Polcari et al. 2024). This event was associated with the very high $SO₂$ emissions witnessed on 2023-05-19 (Figure 33a) which emitted \sim 6 kilotons of SO₂.

Figure 34. Nyamulagira (a) SO₂ emissions and (b) mean INSAR coherence in the caldera during the period 2022/01 – 2024/01. The caldera area inside which the mean coherence is computed (Polcari et al. 2024) is circled in yellow in the geocoded coherence images (top row).

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5.2.6. Thermal activity at Nyiragongo and Nyamulagira volcanoes (DRC) detected by the MIROVA system: 2022-2023 report

In this short report I summarize the satellite thermal observations made by the MIROVA system (www.mirovaweb.it) on the Nyiragongo and Nyamulagira volcanoes during 2022-23.

1. Nyiragongo volcano:

During 2022-2023, the Nyiragongo heat flux was characterized by an extremely variable pattern with periods of low or totally absent thermal emission, alternating with pulsations characterized by VRP > 100 MW. At least 5 thermal pulses have exceeded 1000 MW during the two-years-long analysed period. This occurred on 5 Jan 2022, 24 Mar 2022, 29 Nov 2022, 16 Mar 2023 and 10 Jul 2023. The impulsive thermal activity seems to be indicative of an openvent system (typical of lava lakes) not yet well developed. Low thermal emissions are possibly linked to a continuous degassing on the crater floor, alternating with the sporadic resumption of eruptive activity (Strombolian activity? small lava lakes?) which however is not persistent over time. Overall, approximately 1.9×10^{15} J were issued between 2022 and 2023, with a timeaveraged VRPmean=29.4 MW.

Figure 35. (a) Volcanic Radiative Power (VRP) recorded at Nyiragongo volcano by the MIROVA system during 2022- 2023. (b) Cumulative Radiant Energy recorded in the same period.

2. Nyamulagira volcano:

In the period 2022-2023, the thermal activity of Nyamuragira was continuous and persistent, oscillating between moderate (10 MW < VRP < 100 MW) and very high VRP (1000 MW < VRP < 10000 MW). The phase of greatest radiant flux occurred on 19 May 2023 (23:06 UTC) when a VRP=10.6 GW were reached. This phase occurred in correspondence with a lava "overflow" outside the summit caldera extending about 3-4 km on the west flank and fed by an effusion rate of about 30 m3/s (pers. communication to C. Balagizi). The effusive activity vanished in the following hours and emitted approximately ~0.9 \times 10¹⁵ J. The effusive phase is followed by a period of low thermal activity ending on September 2023. Since then the thermal flux has returned to pre-overflow values. Overall, 4.5×10^{15} J were emitted between 2022 and 2023, with a time-averaged VRP_{mean}= 76.4 MW.

Nyamuragira - Volcanic Radiative Power 10^{10} Very High $10⁹$ High /RP (Watt) 10 Moderate 10 10 $10w$ $\frac{10^5}{\text{Jan }2022}$ **Jul 2022** Jan 2023 **Jul 2023** Jan 2024 $\times 10^{15}$ Nyamuragira - cumulative thermal energy VRPmean=76.4133MW $\overline{4}$ VRE (Joule) 3 $\overline{2}$ ĭ \mathcal{L} Jul 2022 Jan 2023 Jul 2023 Jan 2024

Figure 37. (a) Volcanic Radiative Power (VRP) recorded at Nyamuragira volcano by the MIROVA system during 2021. (b) Cumulative Radiant Energy recorded in the same period.

Publications

Peer reviewed journal articles *Sadiki, A.T., Kyambikwa, A.M., Namogo, D.B. et al. Analysis of the Seismicity Recorded before the May 22, 2021 Eruption of Nyiragongo Volcano, Democratic Republic of the Congo. J. Volcanolog. Seismol. 17, 246–257 (2023). https://doi.org/10.1134/S0742046323700136 Borges D.E. et al. (2023), "Earth observations into action: the systemic integration of earth observation applications into national risk reduction decision structures". Journal of Disaster Prevention and anagement, https://doi.org/10.1108/DPM-09-2022-0186 Cornelis Schwenk, Sophie Negele, Charles M. Balagizi, Werner Aeschbach, Bertram Boehrer (2022) High temperature noble gas thermometry in Lake Kivu, East Africa, Science of The Total Environment, https://doi.org/10.1016/j.scitotenv.2022.155859.*

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- *Georges Mavonga Tuluka (Author), Thystere Matondo Bantidi (Author), Arsène Tumaini Sadiki (2023), Volcanic Seismicity for the East Democratic Republic of Congo: Volcanic Hazard for the Eastern DR Congo. LAP LAMBERT Academic Publishing. ISBN-13 : 978-6206159636.*
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Conference presentations/proceedings

- *M. Polcari, C. Balagizi, E. Ferrentino, S. Valado, D. Coppola, S. Salvi, 2023. Analysis of the lava lake and magma flows during 2023 Nyamulagira volcanic activity detected by SAR coherence variations. In Multi-scale and multi-parametric geodetic monitoring and modeling for studying and forecasting volcanic activity session, COV12, Guatemala*
- *Balagizi, C., Mashagiro, N., Kasereka, M., and Kelly, P.: Pre- and post-eruptive gas composition measurements at Nyiragongo Volcano, East Africa, using a portable Multi-GAS, EGU General Assembly 2023, Vienna, Austria, 24–28 Apr 2023, EGU23-16823, https://doi.org/10.5194/egusphere-egu23-16823, 2023.*
- *Charles M. Balagizi, Ganci, G., Trasatti, E., Tolomei, C., and Beccaro, L.: The 2021 Nyiragongo (DR Congo) eruptive crisis monitored by multi-sensor satellite remote sensing data, EGU General Assembly 2022, Vienna, Austria, 23–27 May 2022, EGU22-12586, https://doi.org/10.5194/egusphere-egu22-12586, 2022.*
- *Balagizi C, 2023. Volcanology in the Democratic Republic of the Congo, Focus on Volcano Monitoring and Volcano Hazards in the Virunga Volcanic Province. 1st International Workshop for African Volcanologists July 24-27, 2023; Nyeri-Kenya*

Research products

The main research products of the Virunga Supersite for the period from January 2020 to December 2021 are included in the present report. Thus, they will be accessible via the link to the report which will be available once the report has been approved; the report will also be available at the GEO GNSL website, Virunga Supersite section at http://geogsnl.org/supersites/permanent-supersites/virunga-supersite/ . However, there is a plan to open a website for the Virunga Supersite which will help to share the research products.

Research product issues

All the research products are not open to the public; some are still under embargo to allow the scientific team to publish their results. However, the results are fully open for hazards assessment and crisis managements, but can be accessed on request.

6. Dissemination and outreach

- Scientific products were already sent to local institutions that are involved in hazards assessment, particularly the GVO that monitors the Nyiragongo and Nyamulagira active volcanoes.
- **The Virunga Supersite has a logo which is shown below:**

Training of local scientists:

- 1°) Albert Kyambikwa, a GVO Junior researcher working in the department of semiology, has finished his two years masters in Volcanology at the University at Buffalo with thesis focusing on the Reconstruction and modeling of lava rheology at Nyiragongo Volcano.
- 2°) Arsène Tumaini Sadiki, a GVO Junior researcher working in the department of semiology, has finished 18 months training in volcano seismology at Istituto Nazionale di Geofisica e Vulcanologia (INGV), Osservatorio Etneo - Sezione di Catania; Italy

- 3°) Olivier Munguiko Munyamahoro, a GVO Junior researcher working in the department of semiology, participated in a short-term scientific training at the Center for the Study of Active Volcanoes; an International Training Course in Volcano Hazards Monitoring. The was held at Hilo campus – University of Hawaii in partnership with the Hawaiian Volcano Observatory, and second part in Vancouver, Washington at the USGS Cascades Volcano Observatory. This training was fully funded by the USGS's Volcano Disaster Assistance Program (VDAP-USAID).
- 4°) Marcellin Kasereka, a GVO Junior researcher working in the department of Geochemistry, is the second year of a master in analytical chemistry focusing on fluids (gases and waters) at the University of Burundi, in Burundi.
- 5°) MOSANGE Thérèse from University of Goma is having a 2 years masters in Environmental Sciences at Abomey Calavi University (Benin), with a scholarship from Abomey Calavi.

7. Funding

The Virunga Supersite has no funding; all the activities were realized by scientists working on a voluntary basis.

8. Stakeholders interaction and societal benefits

The Goma Volcano Observatory (GVO) remains the main and direct beneficiary of the Virunga Supersite services and products. In fact, most of the products are science related and in direct link with volcano study and monitoring. The products are entirely integrated in the routine of the GVO that produces the information to communicate to stakeholders, such as the weekly bulletin about volcano eruptive level, and other communication that is directed to the government and the population.

9. Conclusive remarks and suggestions for improvement

AThe Virunga Geohazards Supersite is presently well established and an operating scientific group formed of local and international scientists and agencies. We expect to get fundings that will allow this motivated team to get instruments for ground-based data collection, local infrastructure development and for the training of the local scientists.

10. Dissemination material for CEOS (discretionary)

Please refer to section 5.